

ESF PESC Exploratory Workshop: How to Constrain the High Density Symmetry Energy Zagreb (Croatia), 15-18 October 2009

Status of the UrQMD predictions

M.Chartier, E.De Filippo, Y.Leifels, R.C.Lemmon, Q.Li, J.Lukasik, A.Pagano, P.Pawlowski, <u>P.Russotto*</u>, W.Trautmann, P.Wu

and the ASY-EOS collaboration



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A key problem: The density dependence of the Symmetry Energy

$$\frac{E}{A}(\rho,I) = \frac{E}{A}(\rho,I=0) + \frac{E_{Sym}}{A}(\rho)I^{2} + \dots O(4) I = \frac{\rho_{n} - \rho_{p}}{\rho_{n} + \rho_{p}}$$

I is the isospin asymmetry in terms of the neutron and protons density

parameterization in the UrQMD



Qingfeng Li et al., J. Phys. G 31 1359-1374 (2005)

parameterization in the IBUU



FIG. 2: (Color online) Symmetry energy as a function of density for the MDI interaction with x = 1, 0, -1 and -2. Taken from Ref. [109].

B.A. Li et al., Phys. Rep. 464 (2008)

Current Experimental Constraints on the Symmetry Energy

Increasingly stringent constraints at $\rho < \rho_0$:

- Giant and Pigmy Dipole Resonances
- Isospin diffusion and n/p ratios in HICs
 Isobaric analogue states and masses
 Isoscaling in projectile fragmentation

PRL 94, 032701 (2005)

PHYSICAL REVIEW LETTERS

week ending 28 JANUARY 2005

Determination of the Stiffness of the Nuclear Symmetry Energy from Isospin Diffusion Lie-Wen Chen,^{1,*} Che Ming Ko,¹ and Bao-An Li²



FIG. 4 (color online). The degree of isospin diffusion as a function of K_{asy} with the MDI and SBKD interactions. γ is the parameter for fitting the corresponding symmetry energy with $E_{sym}(\rho) = 31.6(\rho/\rho_0)^{\gamma}$.

Almost no constraints at $\rho > \rho_0$: • limited π^-/π^+ , K^+/K^0 , n-p flow studies

PRL 102, 062502 (2009) PHYSICAL

PHYSICAL REVIEW LETTERS

week ending 13 FEBRUARY 2009

Circumstantial Evidence for a Soft Nuclear Symmetry Energy at Suprasaturation Densities

Zhigang Xiao,1 Bao-An Li,2.* Lie-Wen Chen,3 Gao-Chan Yong,4 and Ming Zhang1



FIG. 2 (color online). The π^-/π^+ ratio as a function of the neutron/proton ratio of the reaction system at 0.4A GeV with the reduced impact parameter of $b/b_{\text{max}} \leq 0.15$. The inset is the impact parameter dependence of the π^-/π^+ ratio for the 96 Ru + 96 Ru reaction at 0.4A GeV.

From B.Tsang NUFRA2009



URQMD simulations (by Qingfeng Li)



symmetry potential energy.

In order to mimic the strong density dependence of the symmetry potential at high densities, we adopt the form of F(u), used in [4], as

$$F(u) = \begin{cases} F_1 = u^{\gamma}, & \gamma > 0, \\ F_2 = u \cdot \frac{a - u}{a - 1}, & a > 1. \end{cases} \qquad u = \rho/\rho_0 \tag{4}$$

Here, γ is the strength of the density dependence of the symmetry potential. We choose $\gamma = 0.5$ and 1.5, denoted as the symmetry potentials F05 and F15, respectively. *a* (in F_2) is

See Qingfeng Li, J. Phys. G 31 1359-1374 (2005) and references therein





Coalescence condition:

DR <3 fm and DP< 275 MeV/c

(*) W. Reisdorf et al., NPA 612 493-556 (1997)

differential elliptic flow

$$V_2(y, p_t) = \left\langle \frac{p_x^2 - p_y^2}{p_t^2} \right\rangle$$

Q.F. Li and P. Russotto



URQMD simulations: Au+Au @ 400 AMeV

Evolution with impact parameter



FOPI/LAND experiment on neutron squeeze out

Au+Au 400 A MeV



Comparison to experimental data: see W.Trautmann, P.Wu talks

URQMD simulations: Au+Au @ 400 AMeV

5.5<b<7.5 fm



$$V_2(y, p_t) = \left\langle \frac{p_x^2 - p_y^2}{p_t^2} \right\rangle \,.$$

Seen by LAND 37°<θ_{lab}<53° 61°<θ_{lab}<85° 0.3<pt<1.3 GeV/c

inversion of the relative strengths of the elliptic flow for neutrons, protons, and hydrogen \rightarrow relative effect!!!!

URQMD simulations: @ 400 AMeV





Au+Au @ 400 AMeV 5.5<b<7.5 fm





Total

Au+Au @ 400 AMeV 5.5<b<7.5 fm



Au+Au @ 400 AMeV b<2.75 fm



Au+Au @ 400 AMeV 5.5<b<7.5 fm

$$F_{n-p}^{x}(y) \equiv \sum_{i=1}^{N(y)} (p_{i}^{x}w_{i})/N(y),$$
 (28)

where $w_i = 1(-1)$ for neutrons (protons) and N(y) is the total number of free nucleons at rapidity y. Since





Au+Au @ 400 AMeV 5.5<b<7.5 fm



Total

37° < θ_{lab} < **53°**





b<2.75 fm



Double neutron-proton differential transverse flow as a probe for the high density behavior of the nuclear symmetry energy





FIG. 3. (Color online) Rapidity distribution of the isospin asymmetry of free nucleons (upper panels), the difference of the average nucleon transverse flows (middle panels) and the neutron-proton differential transverse flow (lower panels) from $^{132}Sn+^{124}Sn$ reaction at the incident beam energies of 400, 800 MeV/nucleon and b = 5 fm with two symmetry energies of x = 0 and x = -1.

$$F_{n-p}^{x}(y) \equiv \sum_{i=1}^{N(y)} (p_{i}^{x}w_{i})/N(y),$$
 (28)

where $w_i = 1(-1)$ for neutrons (protons) and N(y) is the total number of free nucleons at rapidity y. Since



FIG. 1. (Color online) Density dependence of nuclear symmetry energy using the MDI interaction with x = 0 and x = -1.



FIG. 4. (Color online) Same as the lowest two panels (c) and (f) of Fig. 3 but for the reaction system of ¹¹²Sn+¹¹²Sn.

URQMD simulations: Au+Au @ 400 AMeV b=5.5-7.5 fm



$$F_{n-p}^{x}(y) \equiv \sum_{i=1}^{N(y)} (p_{i}^{x}w_{i})/N(y),$$
 (28)

where $w_i = 1(-1)$ for neutrons (protons) and N(y) is the total number of free nucleons at rapidity y. Since

Proton	and
neutron	
All pt	
Slopes	
0.08	
0.08	

URQMD simulations: Au+Au @ 400 AMeV

b=5.5-7.5 fm



$$F_{n-p}^{x}(y) \equiv \sum_{i=1}^{N(y)} (p_{i}^{x}w_{i})/N(y),$$
 (28)

where $w_i = 1(-1)$ for neutrons (protons) and N(y) is the total number of free nucleons at rapidity y. Since

Proton and neutron pt>0.75 GeV/c Slopes 0.25 0.18

URQMD simulations: Au+Au @ 400 AMeV

b=5.5-7.5 fm



```
p and neutron
pt>0.75 GeV/c
37° <θ<sub>lab</sub> < 53°
61°<θ<sub>lab</sub><85°
σ<sub>rp</sub>=24°
Slopes
0.24
0.16
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400 AMeV (SMDBHF)

¹²⁴Sn+¹²⁴Sn b=4-6 fm

¹¹²Sn+¹¹²Sn b=4-6 fm



 \rightarrow DFpn and so on....

From B.Tsang NUFRA2009



Symmetry Energy Project →International collaboration to determine the symmetry energy over a range of density Require: New Detectors (TPC), travel money, theory support

R.Lemmon, P.Russotto et al. ASY-EOS experiment approved by GSI-PAC (possible) 1st phase toward FAIR ??? (e.g. ¹³²Sn,¹⁰⁶Sn beams)

Au+Au @ 400 AMeV ⁹⁶Zr+⁹⁶Zr @ 400 AMeV ⁹⁶Ru+⁹⁶Ru @ 400 AMeV

Other detectors in order to determine reaction plane or measure t,³He, N/Z of light IMFs....?????



Figure 3: Schematic diagram of experimental setup in Cave C.

Experiment setup (not in scale)

ALADIN-CHIMERA

500 cm 210 cm 120 cm DARK!!!! θ=**<24° θ>=7°** θ<6.28° target 16X16 cm² hole beam Aladin Tof **CHIMERA** Wall 448 CsI Charged fragments p,H,n * $\theta_{lab} = 45^{\circ} \pm 8^{\circ}$ LAND

Artistic view (not in scale)

black→all charged particles

blue-red→used in reaction plane reconstrution (Q vector)



We define a vector constructed from the transverse momenta p_{ν}^{\perp} of detected particles:

$$\boldsymbol{Q} = \sum_{\nu=1}^{M} \omega_{\nu} \boldsymbol{p}_{\nu}^{\perp}, \qquad (1)$$

 $\omega_{\nu} = 1$ for $y_{\nu} > y_{c} + \delta$, $\omega_{\nu} = -1$ for $y_{\nu} < y_{c} - \delta$, and $\omega_{\nu} = 0$ otherwise. For symmetric collisions it is nat-

Using only θ , ϕ , Z information $\vec{Q} = \sum_{\nu=1}^{M} \omega_{\nu} Z_{\nu}^{\perp}$ Only y_{cm}>0.1



CHIMERA ∆t (Csi) =10 ns ∆E/E (CsI) =10 %









Pulse-shape G4I

Gates (stretched and slow) = 50 ns (QDC V792, 50 Ohm impedence input / 400pC range) Standard preamplifier 45 mV/MeV



0.5 cm plastic scintillator (phototube readout) in front of G4I CsI(TI). Time-of-Flight respect to the beam start detector



Useful to determine CsI average time resolution

Estimeted Csl time resolution ~ 8-11 ns.

G2-E – G2-I

Gates (stretched and slow) = 500 ns (QDC V792, 1.1 kOhm impedence input*) Standard preamplifier 45 mV/MeV (Standard Si-Csi Chimera telescope)



DPSA: almost all Si and CsI signals splitted for digital analysis







Conclusions

Several heavy Ion reactions observables have been proposed in order to get information on symmetry energy (giant and pigmy dipole resonances, isobaric analogue states and masses, isospin diffusion, π⁻/π⁺, K⁺/K⁰, n/p ratios, neutron/proton differential flow, v2, ³H/³He ratio...).

More extended data sets and consistency checks are needed in order to arrive at firm conclusions especially at supra-saturation densities

Even if there are a lot of open questions (N-N cross section, effective masses,...) in the ASY-EOS experiment at GSI we will try to measure crucial observables......

....."a good constraint" of symmetry energy at supra-saturation density ???

The high density behaviour of the nuclear symmetry energy $E_{sym}(\rho)$ is very important for understanding both high density nuclear matter and many interesting astrophysical objects, but it is also subject to the large uncertainty.